

**Erratum: GW170104: Observation of a 50-Solar-Mass Binary
Black Hole Coalescence at Redshift 0.2
[Phys. Rev. Lett. **118**, 221101 (2017)]**

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This Erratum reports an error found in the implementation of the code of the LIGO Scientific and Virgo Collaborations (LVC) as used in gravitational-wave-based estimations of possible deviations from the post-Newtonian (PN) terms expected in general relativity (GR). The error concerned the 0.5 PN term and affected the results previously published for GW150914 [1] in Ref. [2], for GW151226 [3] in Ref. [4], and for GW170104 [5]. We corrected the bug and present the reproduced results in this Erratum, as well as in the related Errata [6,7]. The main conclusion, that the results are consistent with general relativity, remains.

The test for the parametrized post-Newtonian [8] deviations from the expected GR values relied on creating non-GR waveforms [2,9–13] and using them as potential matches for the observed waveforms [14–17]. In these waveforms, implemented in the frequency domain, freedom was introduced by allowing the phase coefficients describing different powers of the post-Newtonian parameter (equivalently, powers of the frequency) to assume a range of values, not only the particular values prescribed by GR.

However, a coding bug was introduced, identically zeroing the deviations at 0.5 PN in the inspiral regime (as in GR). The 0.5 PN deviations were hence absent in the phasing formula, though not in the junction conditions that relate the inspiral regime to the intermediate regime. Any constraints obtained in [2,4,5] only resulted from the latter.

This error affected the results of the non-GR parameter estimation (PE) [14] pipeline tests performed for finding bounds on possible PN deviations from GR. In particular, they affect the bounds on the single deviations in the 0.5 PN term and on the tests with multiple deviations together. These erroneous results appeared in Figs. 6 and 7 and Table I of [2], in Figs. 7 and 8 of [4], and in Fig. 9 of the Supplemental Material of [5]. The corrected versions of all of these have been produced.

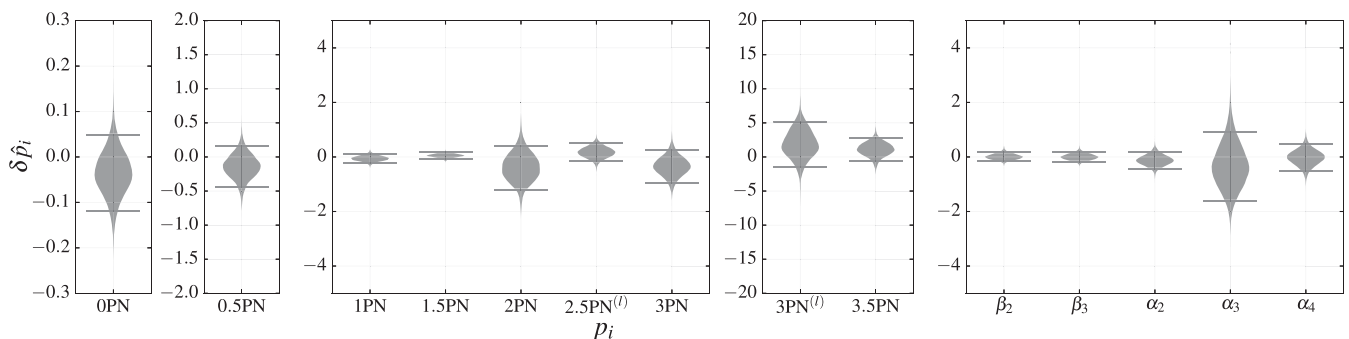


FIG. 1. This is the corrected Fig. 9 in the Supplemental Material of [5], showing the violin plots for the parametrized test, combining posteriors for GW170104 with the two confident detections made in LIGO’s first observing run, GW150914 and GW151226.

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The correction for the Supplemental Material of [5] appears below, while the others are available in [6,7]. All these results are consistent with GR.

We also correct here a misprint in Eq. 5 of the Supplemental Material: the correct equation is,

$$p(\theta|e_1, \dots, e_N) \propto p(\theta) \prod_i^N p(e_i|\theta).$$

The error, introduced by erroneous caching during the optimization of the waveform generation for efficient PE, has been corrected in commit [18] of the LALSuite [19] code. No subsequent LVC papers have been affected.

Note that, while this error also affected the analysis of GW170608 [20], the reported results require no changes: with the corrected analysis, the GR-predicted PN coefficient values continue to be consistent with the data. No change is required regarding the preliminary reported results for GW170814 [21] either.

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